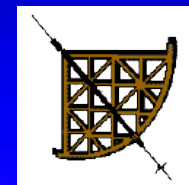# Bologna Open Clusters Chemical Evolution project

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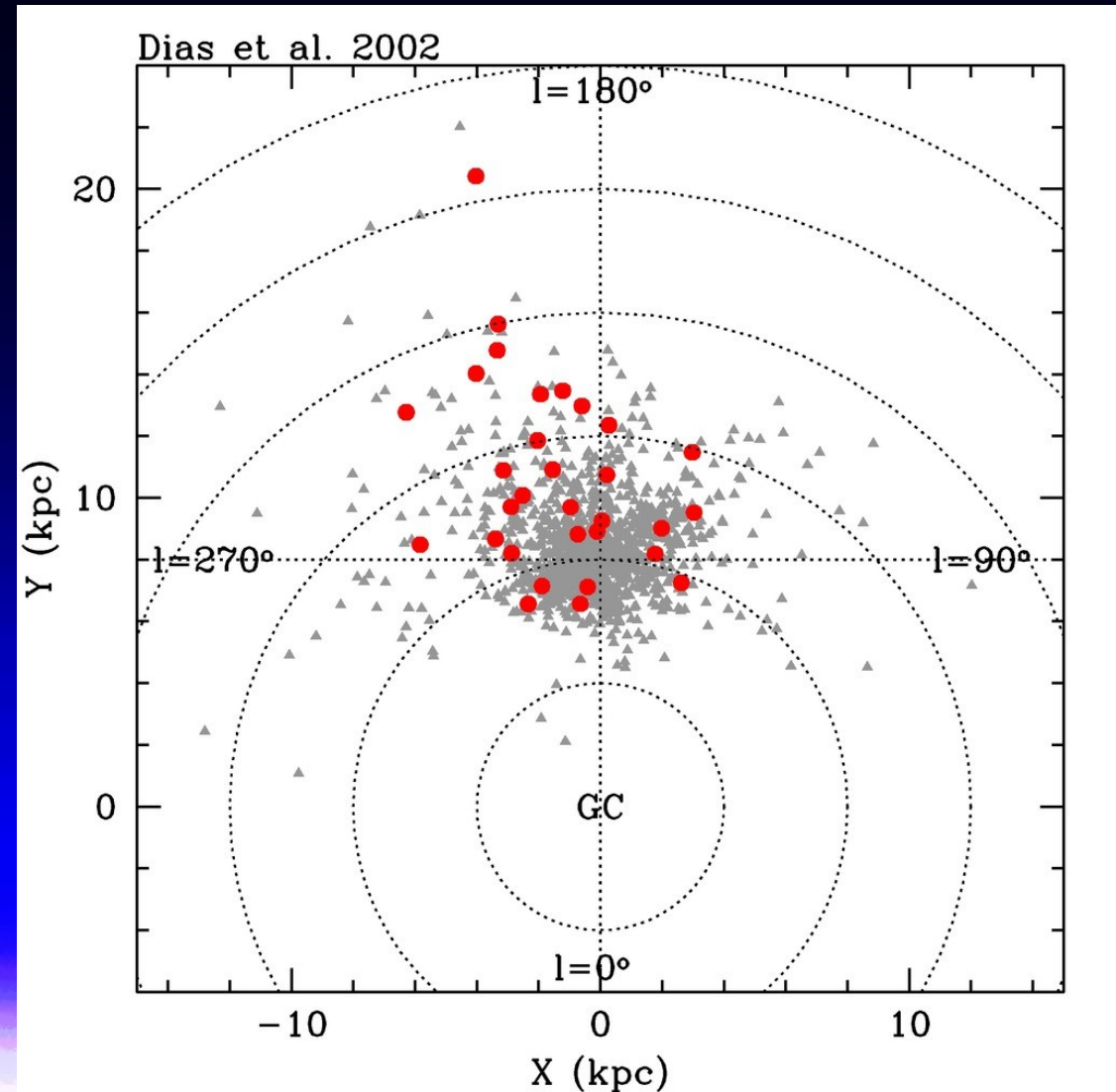
BOCCE members

Monica Tosi
Angela Bragaglia

- Andrea V. Ahumada
- Gloria Andreuzzi
- Giacomo Beccari
- Michele Cignoni

- Eugenio Carretta
- Raffaele Gratton
- Sarunas Mikolaitis
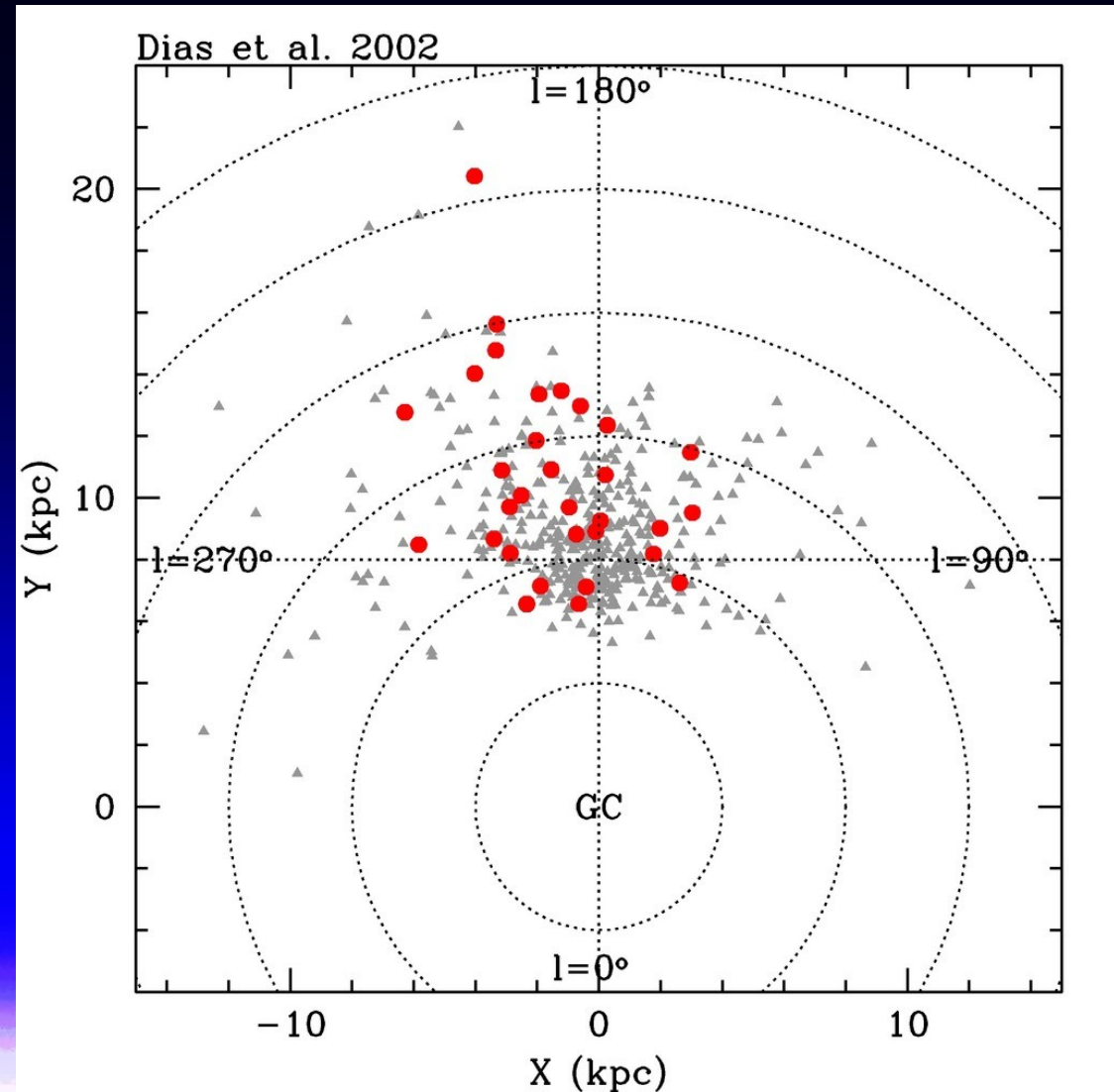- Grazina Tautvaisiene

...and many more!

Observing OCs

- Why

- How many objects

- What we observe
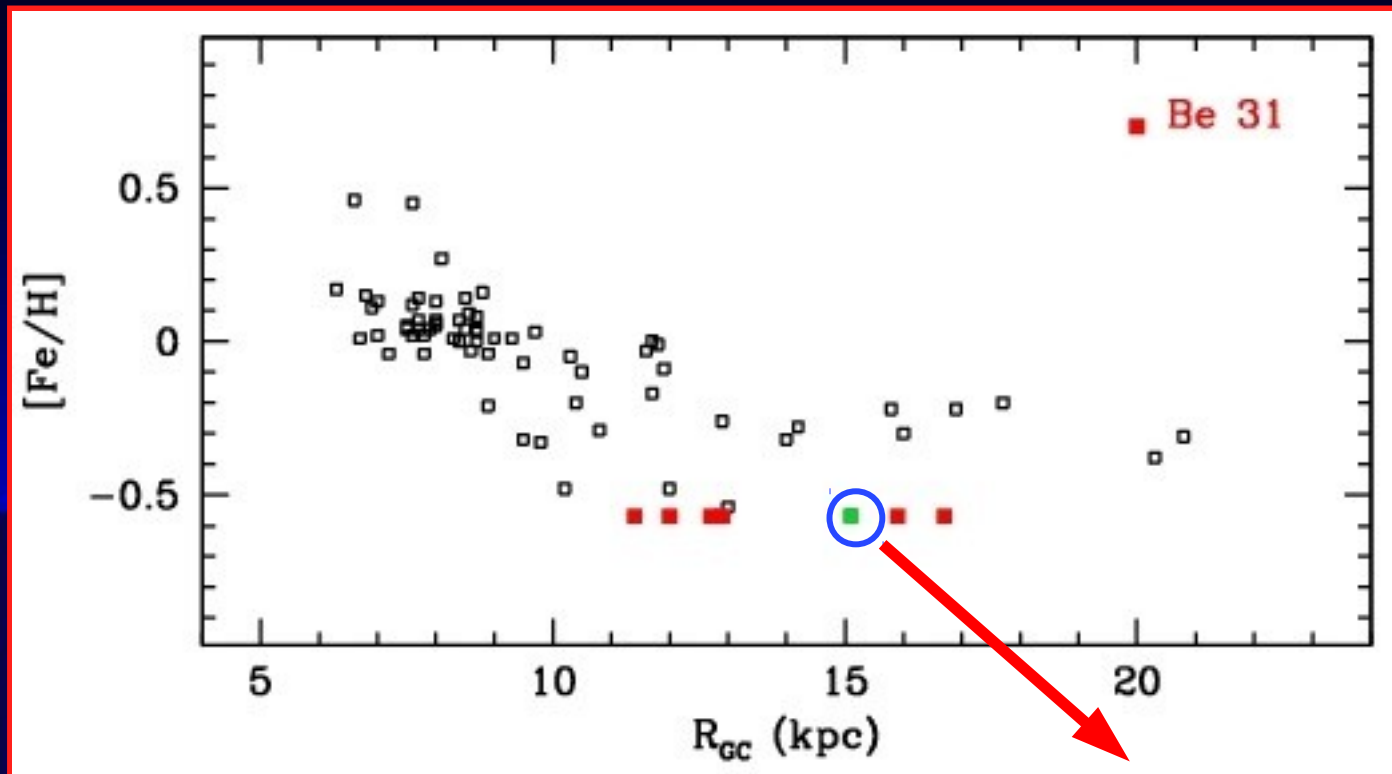
- What we measure

- GAIA

Observing OCs

- Why

- How many objects

- What we observe

- What we measure

- GAIA

Precision and Accuracy

• E.g.: Berkeley 31

Found in literature:
4 kpc < d < 8 kpc
2 Gyr < age < 8 Gyr
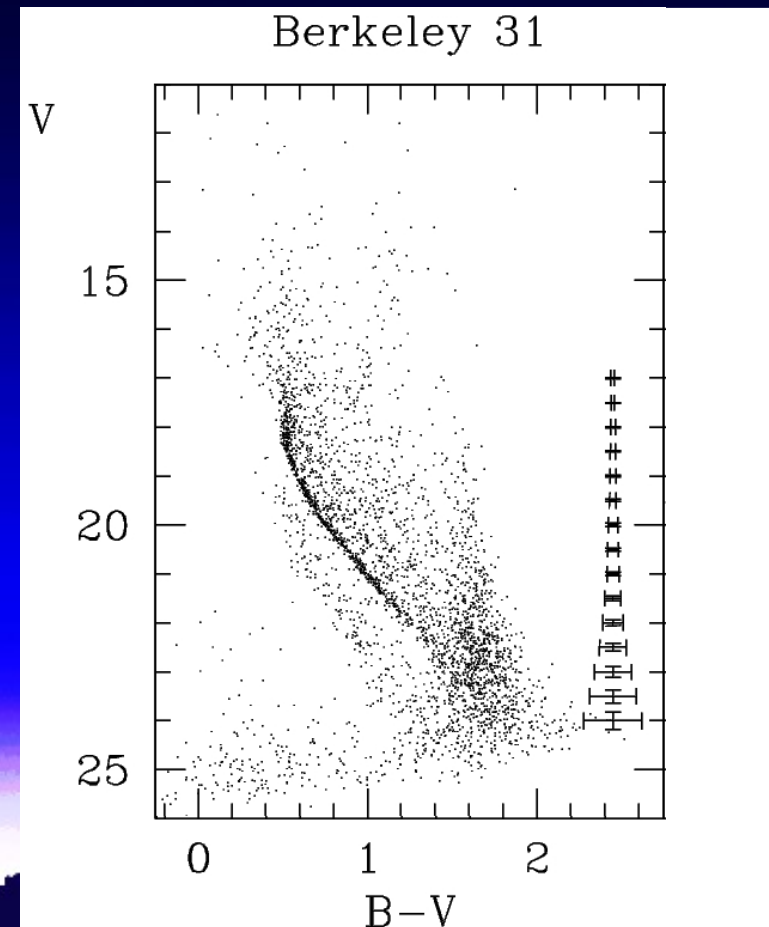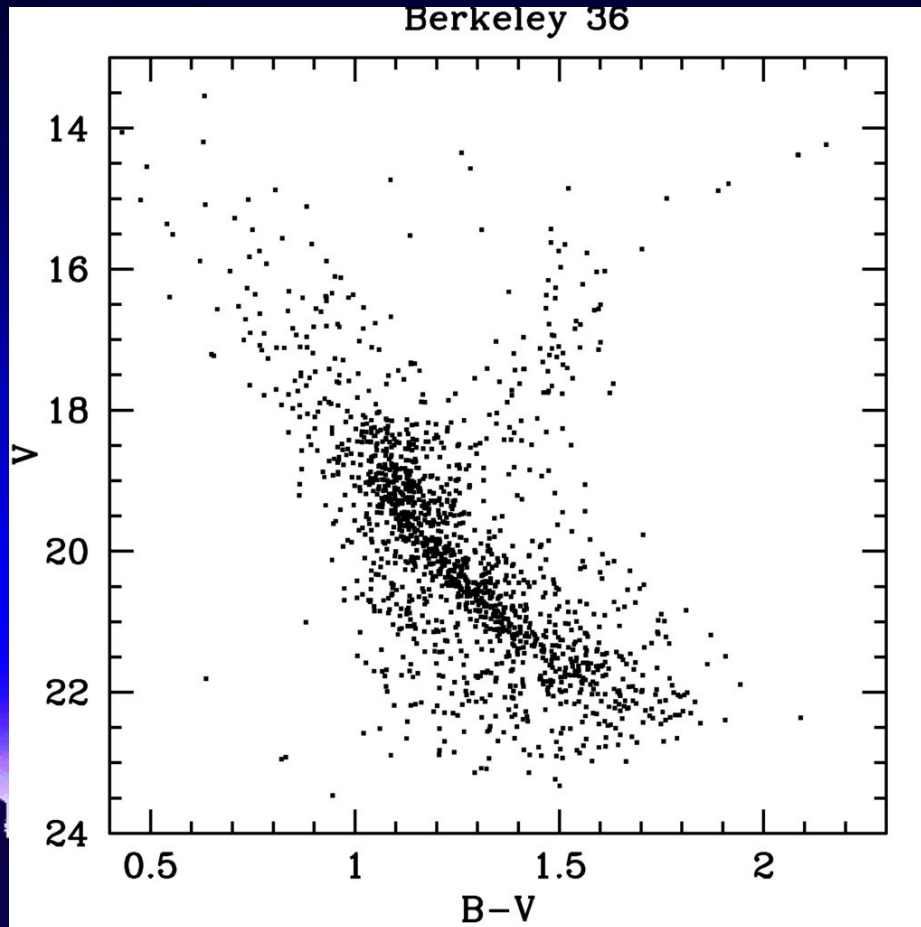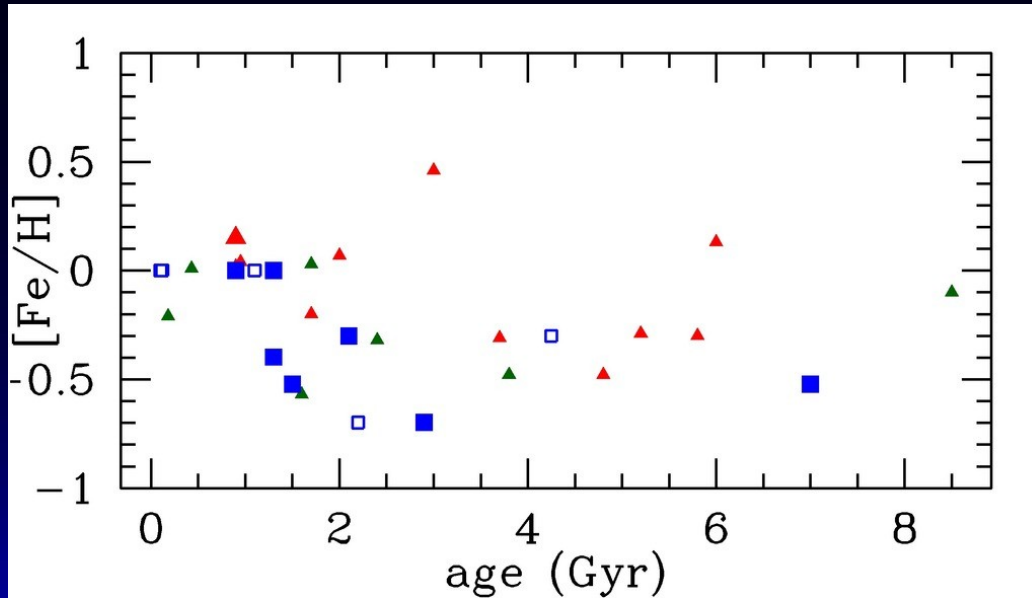
Our estimate:
2.3 Gyr < age < 2.5 Gyr
Rgc > 15 kpc

Brand New Clusters (photo)

- Be 23, Be 31, King 8

- Be 27, Be 34, Be 36
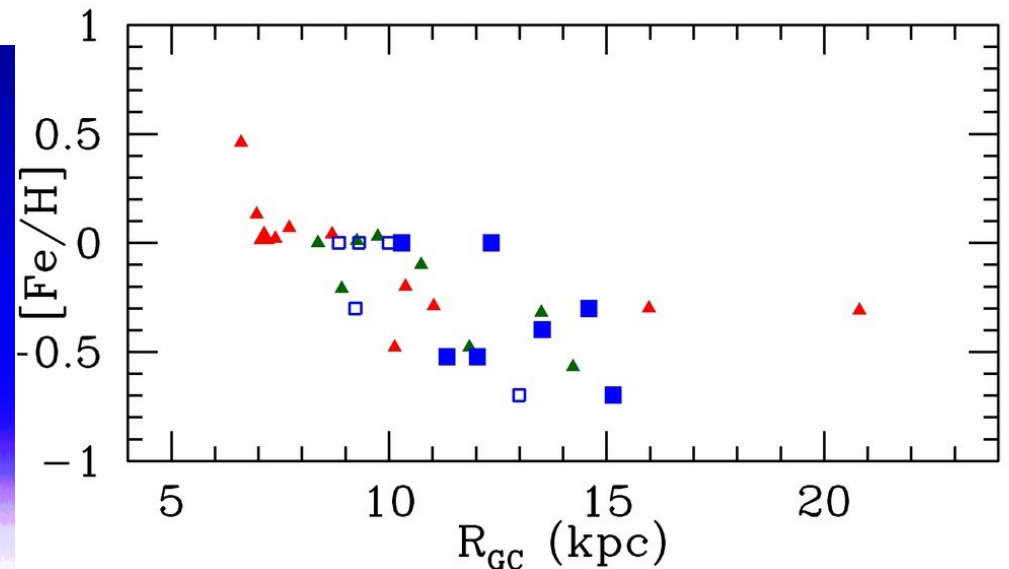
- NGC 2849, NGC 6134

- and more with LBC@LBT

Spectroscopic data

- Δ[Fe/H] ~ 0.1 dex
- About 2/3 (20 OCs) has HiRes spectroscopy
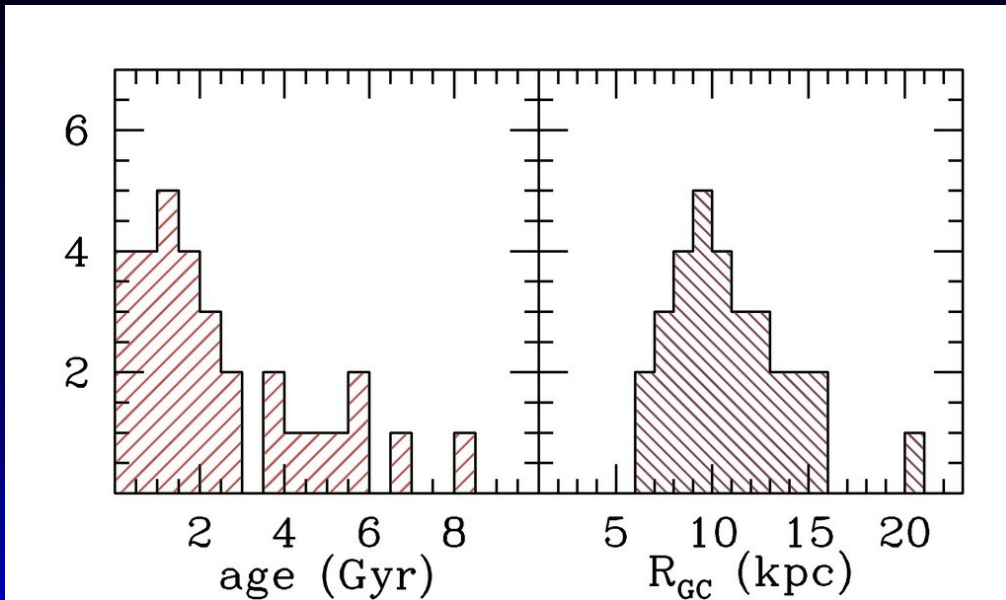- Only ~ 80 OCs has HiRes spectroscopy in DAML02

The Gaia Eso Spectroscopic Survey will target 10^5 stars and 100 OCs

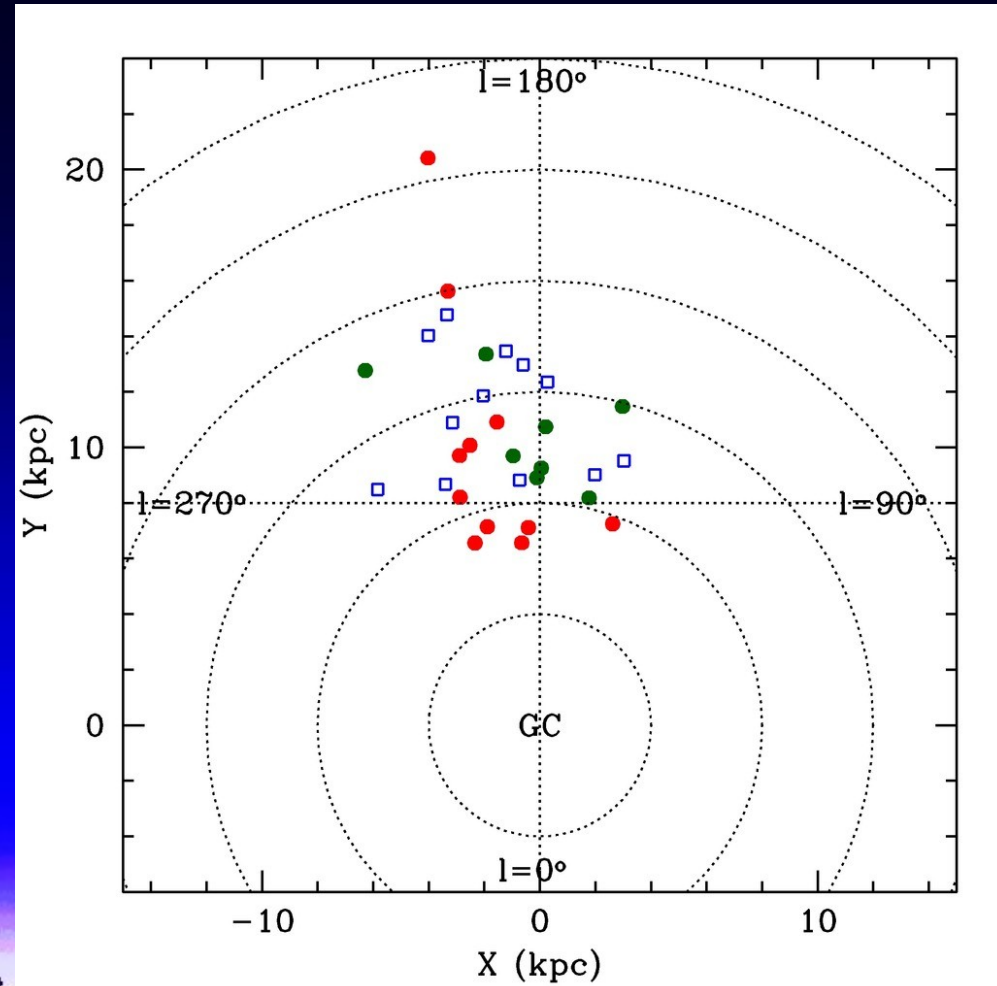We have photometry for many clusters and on large field of view

BOCCE

Spatial distribution of the BOCCE clusters

Age and distance distribution of the OCs in the BOCCE database

Thank you!