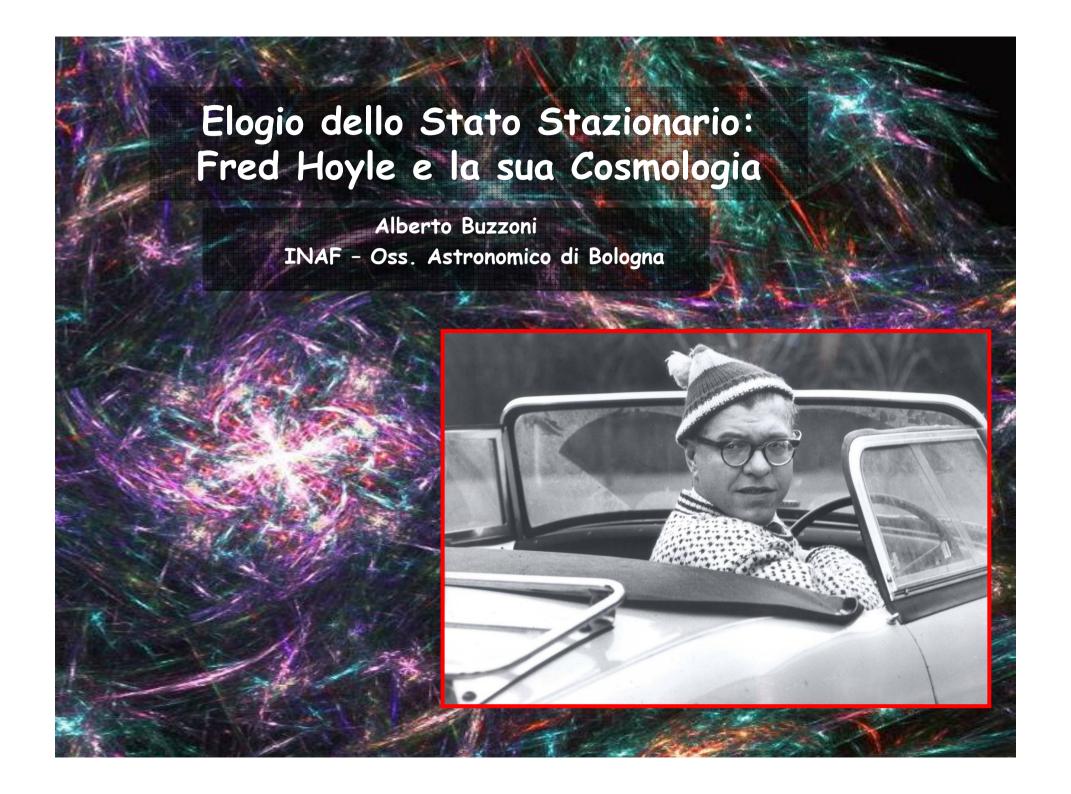
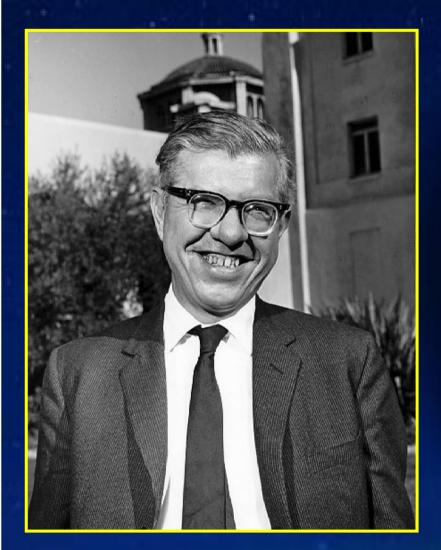
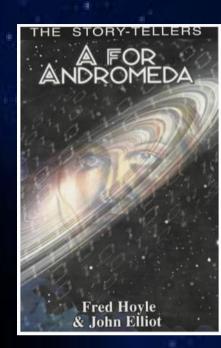
alla difficoltà di conciliare la sua teoria con le leggi di conservazione della fisica, Hoyle replicava:

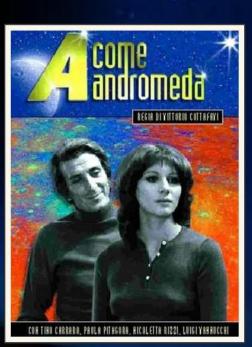
"Forse è paradossale. Ma non è ancora più paradossale l'idea che un bel sacco di roba, l'intero Universo, sia nato, in un attimo, dal niente?"

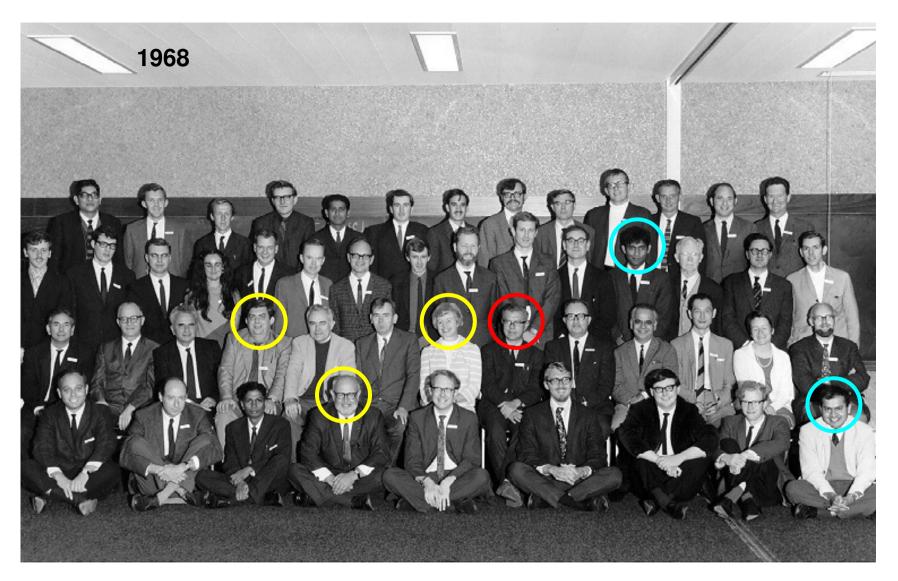












The Institute of Theoretical Astronomy summer conference in 1968 was on cosmic rays, and Hoyle drew many famous physicists to it. July 1968 in the IOTA conference room, Cambridge. Among the many famous faces are: Row 1: Ramaty (left); Meyer (2); Fowler (4); Clayton (6); Narlikar (right). Row 2: Sciama (left); Shapiro (2); Peters (3); Burbidge (4); Simpson (5), P. Fowler (6); Burbidge; Hoyle; McCrea(9); Hayakawa (11). Row 3: Fichtel (3rd); Trimble (4); Walker (5); Parker (6); Taylor (9); Pagel (11); Wickramasinghe (12); Wolfendale (14). Top Row: Price (2); Reeves (3); Rees (6); Arnett (8); Sargent (9); Wagoner (10); Bolton (11); Solomon (12).

5 October 1956, Volume 124, Number 3223

SCIENCE

B²FH (1956)

Origin of the Elements in Stars

F. Hoyle, William A. Fowler, G. R. Burbidge, E. M. Burbidge

Experimental (1, 1a) and observational (2-6) evidence has continued to accumulate in recent years in support of the theory (7-10) that the elements have been and are still being synthesized in stars. Since the appearance of a new and remarkable analysis by Suess and Urey (11) of the abundances of the elements, we have found it possible to explain, in a general way, the abundances of practically all the isotopes of the elements from hydrogen through uranium by synthesis in stars and supernovae. In this article we wish to outline in a qualitative fashion the essentially separate mechanisms which are required in stellar synthesis

Thermal Conversion of Pure Hydrogen through Helium to Iron

As long as extremely high temperatures in excess of 5×10^9 degrees Kelvin are not under consideration, the general tendency of nuclear reactions inside stars to increase the average binding energy per nucleon. For a given temperature and density and for a given time scale of op-

in some cases by resonance penetration. Since barrier effects become less severe as the temperature increases, it follows that the binding energies increase with temperature. This will become clear from

At temperatures from about 107 to

 5×10^7 degrees in main-sequence stars, hydrogen is transformed to helium, 4H1 → He4, with an average binding energy of 7.07 million electron volts (Mev) per nucleon. We emphasize that the proton-proton sequence of reactions makes possible the production of helium starting only with hydrogen, The recent discovery of the free neutrino as reported by Cowan et at. (1a) leads to increased confidence in the existence of the primary proton-proton interaction which proceeds through prompt electron-neutrino emission. At temperatures from 108 to 2 × 108 degrees in red giant stars, He4 is transformed principally to C12, O16, and Ne20 with an average binding energy of 7.98 Mev per nucleon. The important roles of the ground state of Be8 and of the second excited state of C12 in expediting the primary process of helium

fusion, 3He4 -> C12, have recently been

appreciably greater atomic weight than Fe⁵⁶.

The situation, then, is that a thermal "cooking" of pure hydrogen yields principally He⁴ and the α -particle nuclei with A = 4n, Z = 2n, n = 3. 4. 5. 6. 7. 8. 9. and

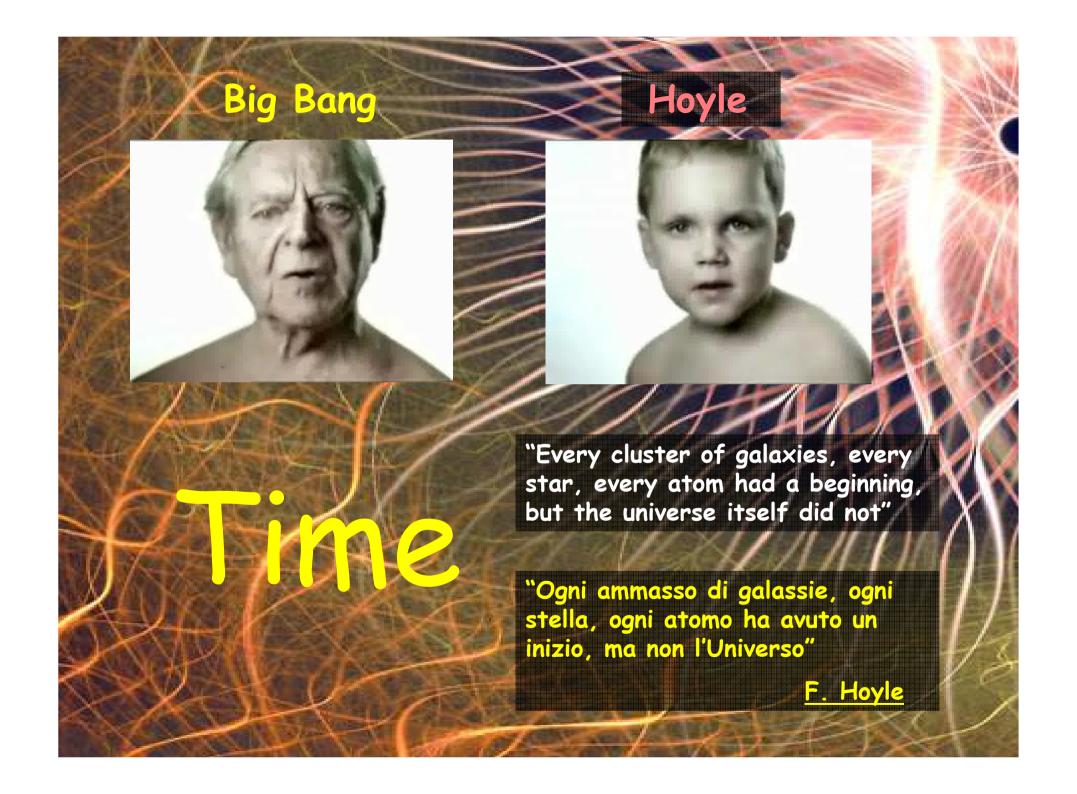
10 (C12 to Ca40), to centered around Fe56. abundant nuclei. Mor abundances that have for these nuclei, and 20-odd isotopes of tit chromium, manganese nickel, show good ag served abundances. Th rium calculations by been considerably im into account the lowof the iron-group nucle active nuclei which u them, and by statistica state according to its that expected on nu Typical results for the dances of the chron 3.8 × 109 degrees are in

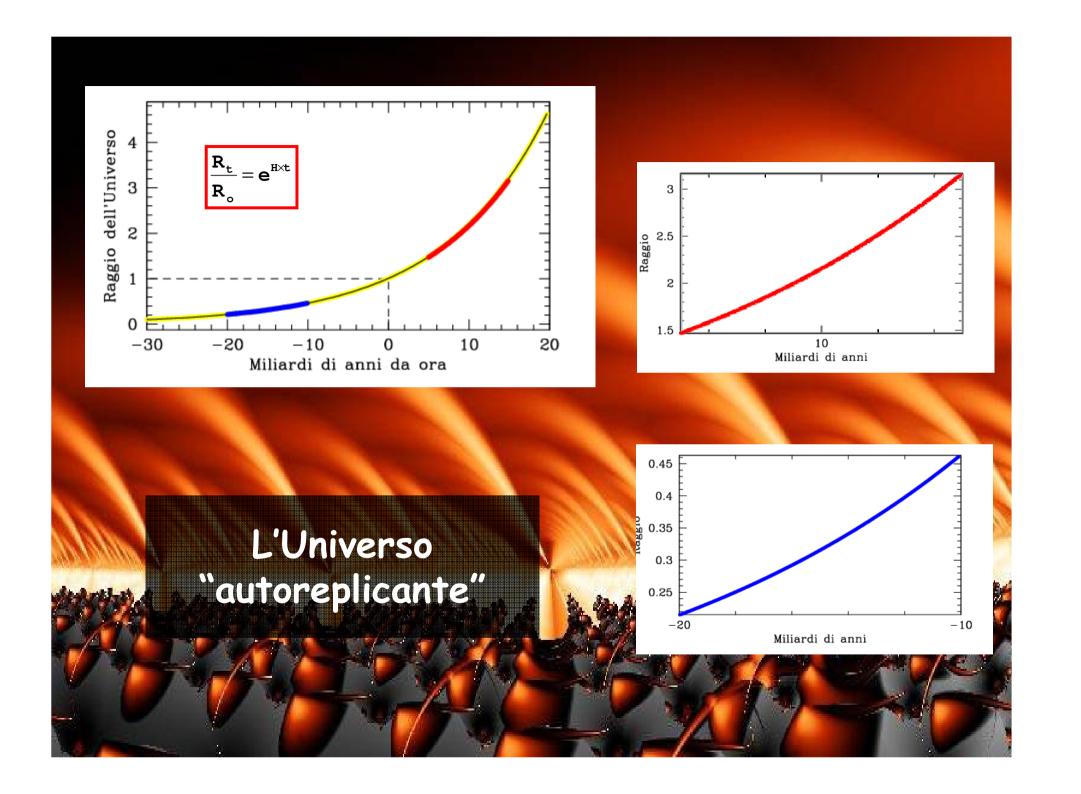
We regard results si sented in Table 1 as giv to the view that the el sideration were synth and that they became tributed in space, eithe from late-type giants explosion, as for insta

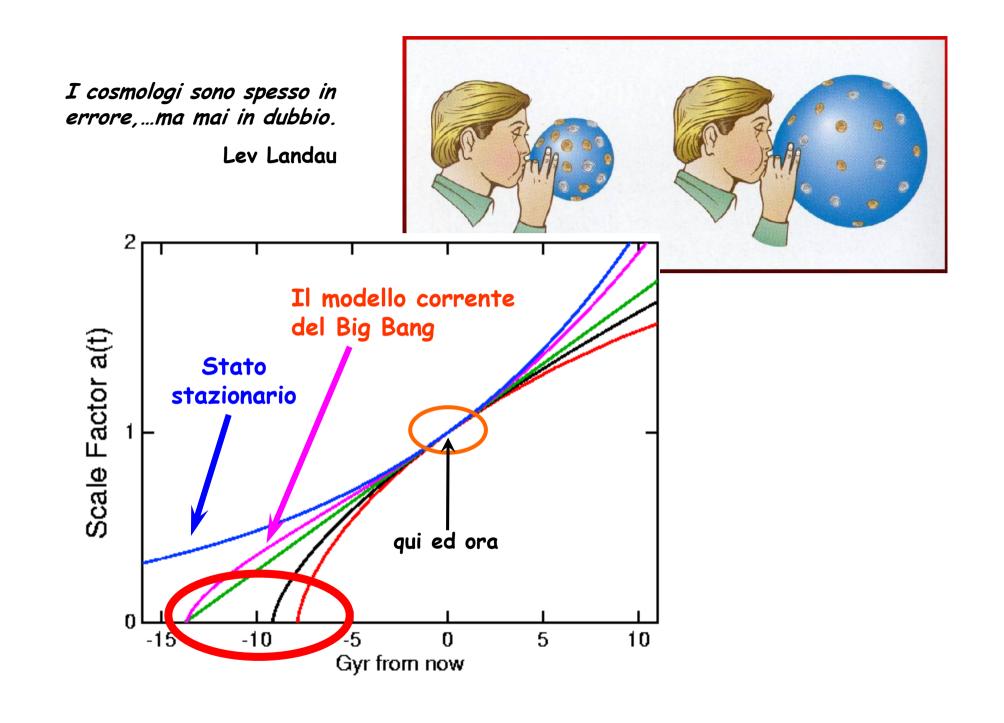
Thermal Reactions of and Helium with Light Institute of Physics

More complicated enects arise when the thermal cooking is considered, not of completely pure hydrogen, but of hydrogen adulterated with a small proportion of the elements mentioned in the previous paragraphs. When a second-genera-









Hubble Diagram 40,000 veloce lontano 20,000 Circa 2 miliardi di anni luce Noi, qui, ora Distance to galaxy (Mpc)

H e' costante?

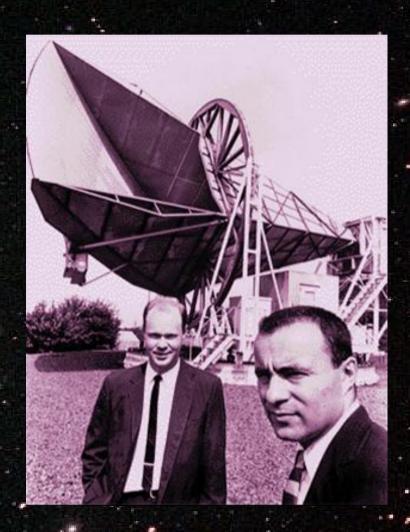
Hoyle: Si!

"gli altri": Nooo!

Il problema della metrica



La radiazione di fondo (1964)



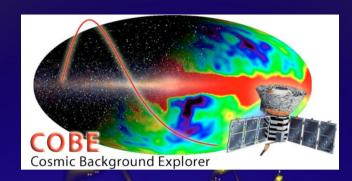
"Philosophically, I liked the steady-state cosmology. So I thought that we should report our results as a simple measurement; the measurement might be true after the cosmology was no longer true!"

"Dal punto di vista filosofico, io ho sempre amato la cosmologia dello stato stazionario. Per questo ho sempre pensato che dovremmo considerare il nostro risultato semplicemente come una misura: dopotutto, una misura può essere corretta anche se la cosmologia finisse per non esserlo!"

R.W. Wilson (1990)



COBE Cosmic Background Explorer (1989)



FIRAS data with 400o errorbars

15

2.725 K Blackbody

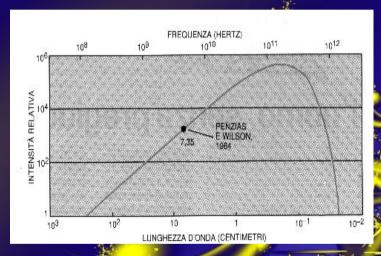
0.5

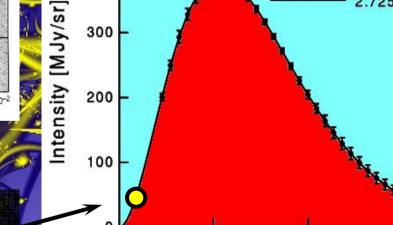
20

Wavelength [mm]

10

ν [/cm]





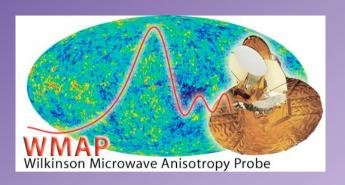
400

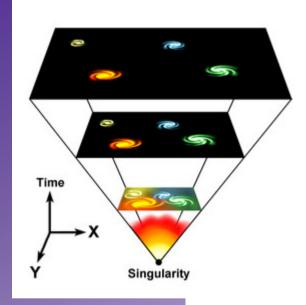
300

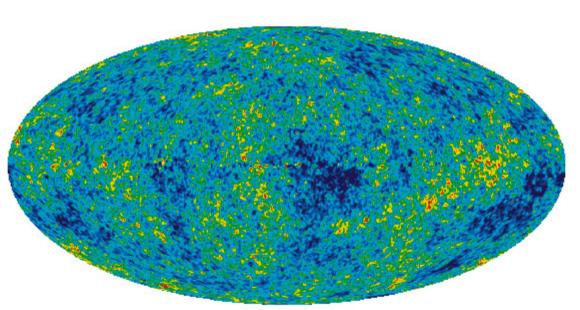
200

Penzias & Wilson

WMAP Wilkinson Microwave Anisotropy Probe (2001)

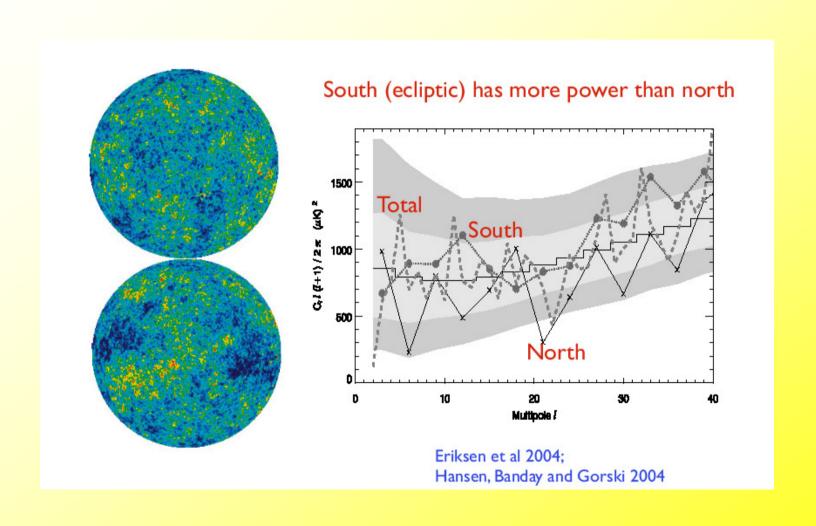




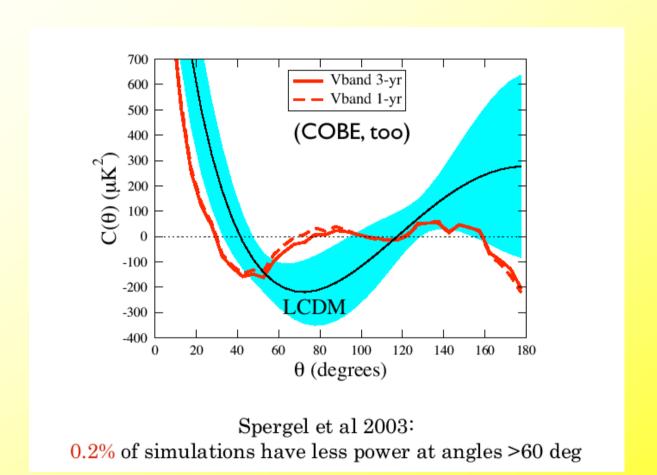


ILC map, WMAP collaboration

Strane (a)simmetrie...



Le strutture cosmiche su larga scala



1

PRL 93, 221301 (2004)

PHYSICAL REVIEW LETTERS

week ending 26 NOVEMBER 2004

Is the Low-ℓ Microwave Background Cosmic?

Dominik J. Schwarz, ¹ Glenn D. Starkman, ^{1,2} Dragan Huterer, ² and Craig J. Copi²

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(Received 26 March 2004; revised manuscript received 14 September 2004; published 23 November 2004)

The large-angle (low-ℓ) correlations of the cosmic microwave background exhibit several statistically significant anomalies compared to the standard inflationary cosmology. We show that the quadrupole plane and the three octopole planes are far more aligned than previously thought (99.9% C.L.). Three of these planes are orthogonal to the ecliptic at 99.1% C.L., and the normals to these planes are aligned at 99.6% C.L. with the direction of the cosmological dipole and with the equinoxes. The remaining octopole plane is orthogonal to the supergalactic plane at 99.6% C.L.

DOI: 10.1103/PhysRevLett.93.221301

PACS numbers: 98.70.Vc

Much effort is currently being devoted to examining the cosmic microwave background (CMB) temperature anisotropies measured with the Wilkinson Microwave Anisotropy Probe (WMAP) [1–4] and other CMB experiments [5]. While the data are regarded as a dramatic confirmation of standard inflationary cosmology, anomalies exist. In particular the correlations at large angular separations, or low ℓ , exhibit several peculiarities.

Most prominent among the "low- ℓ anomalies" is the near vanishing of the two-point angular correlation function $C(\theta)$ at angular separations greater than about 60 degrees. This was first measured using the Cosmic Background Explorer's differential microwave radiometer [6] and recently confirmed by observations with WMAP [4]. This anomalous lack of large-angle correlation is connected to the low value of the quadrupole contribution C_2 in a spherical harmonic expansion of the CMB sky—

$$\Delta T(\theta,\phi) \equiv \sum_{\ell=1}^{\infty} T_{\ell} \equiv \sum_{\ell=1}^{\infty} \sum_{m=-\ell}^{\ell} a_{\ell m} Y_{\ell m}(\theta,\phi) \quad (1)$$

[with $(2\ell+1)C_\ell=\sum_m |a_{\ell m}|^2$]—although the smallness of C_2 does not fully account for the shape of $C(\theta)$. The

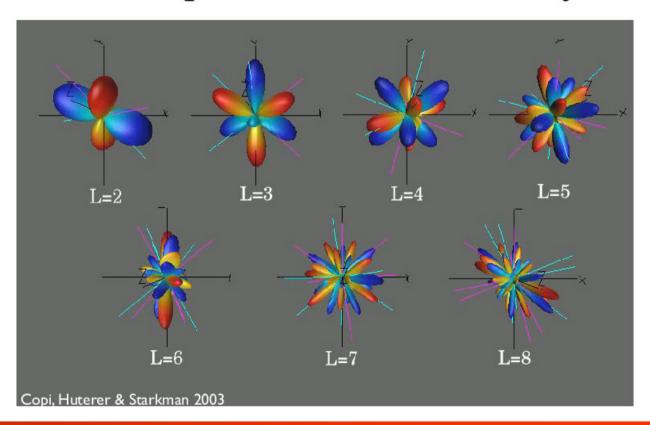
and octopole are maximized are unusually aligned, $|\mathbf{n_2} \cdot \mathbf{n_3}| = 0.9838$. Eriksen *et al.* [11] found that the deficit in large-scale power is due to a systematic deficit in power between $\ell=2$ and 40 in the north ecliptic hemisphere compared to the south ecliptic hemisphere. Some of us [12] have shown that the $\ell=4$ to 8 multipoles exhibit an odd, but very unlikely (~1% probability), correlation with $\ell=2$ and $\ell=3$. These low- ℓ anomalies (and others [13]) have all been pointed out before, but no simple connection has been made between them. Here we remedy that situation.

By far the largest signal in the CMB anisotropy is the dipole, recently measured by WMAP [1] to be (3.346 ± 0.017) mK in the direction $(l = 263.85 \circ \pm 0.1 \circ$, $b = 48.25 \circ \pm 0.04 \circ$) in galactic coordinates. This is caused by the motion of the Sun with respect to the rest frame defined by the CMB. As shown by Peebles and Wilkinson [14], the dipole induced by a velocity v is $\bar{T}(v/c)\cos\theta$, where θ is measured from the direction of motion. Given $\bar{T} = (2.725 \pm 0.002)$ K [15], one infers that $v \approx 370$ km/s.

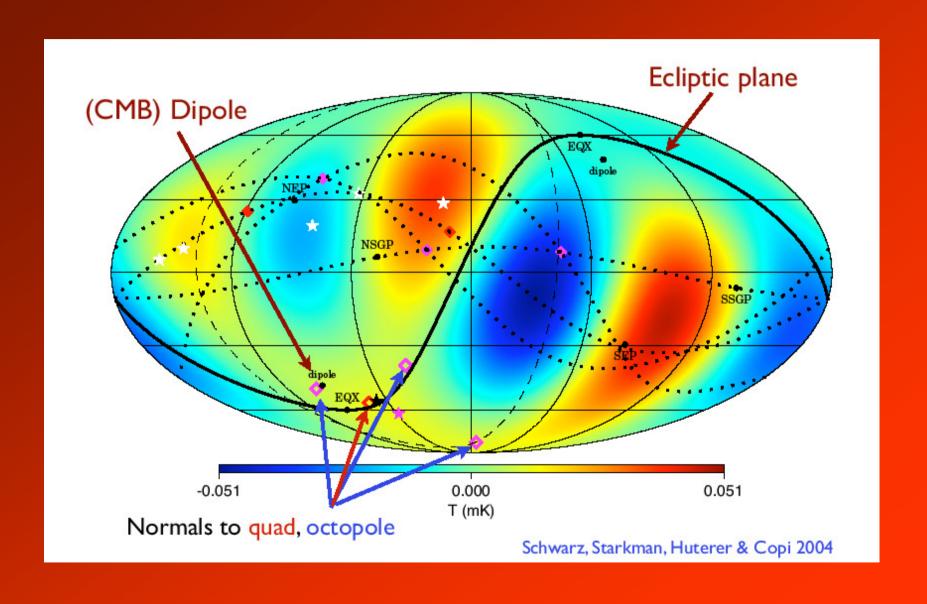
The solar motion also implies [14,16,17] the presence of a kinematically induced Doppler quadrupole (DQ). To second order in $\beta \equiv v/c \approx 10^{-3}$, an observer moving with respect to the CMB rest frame sees the usual mono-

Il problema dei Multipoli e il loro orientamento

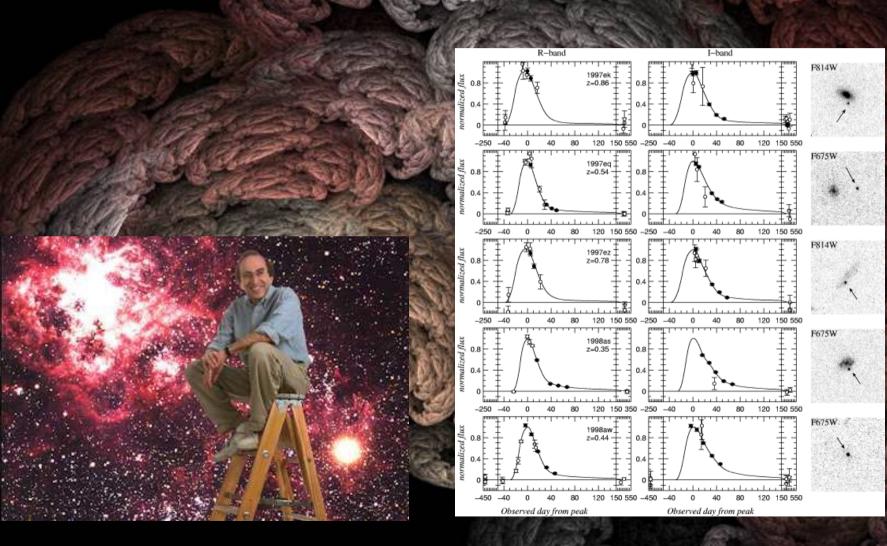
Multipole vectors of our sky

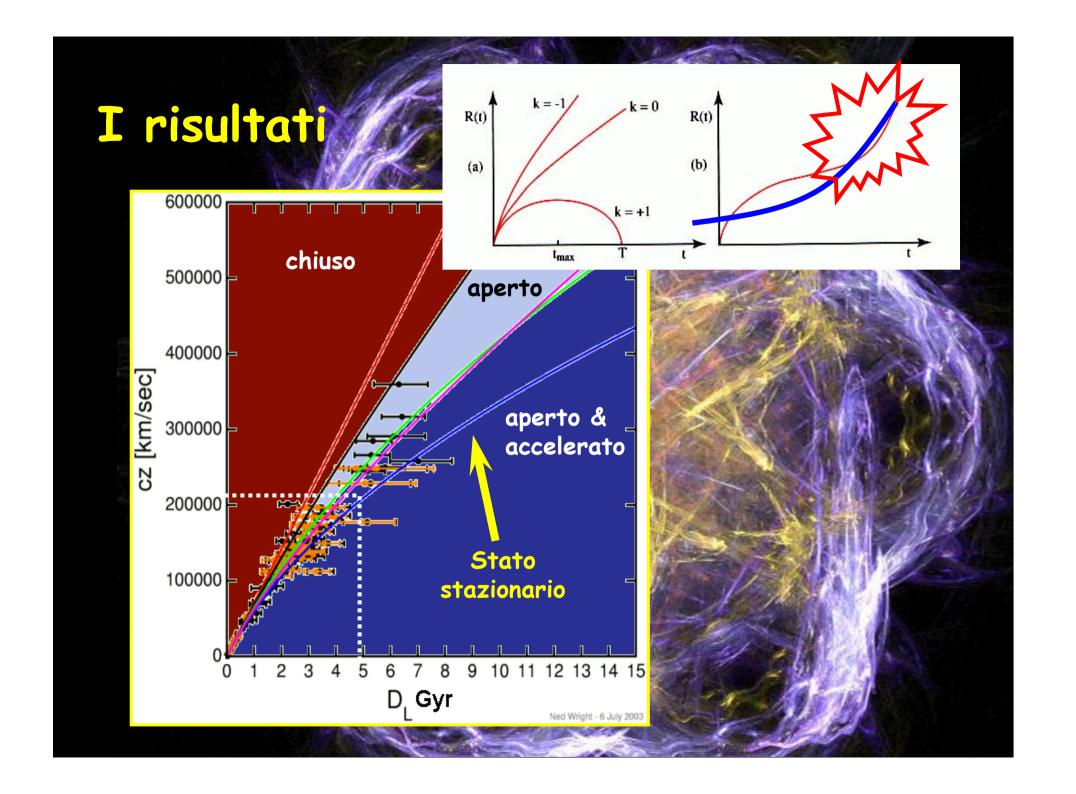


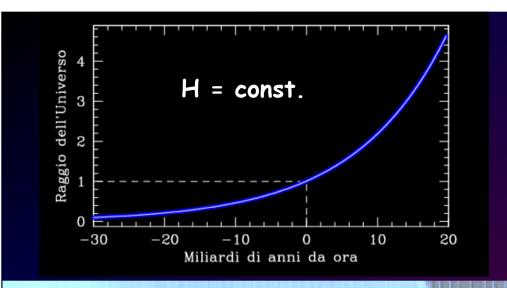
Congiura "cosmica"?



Le Supernovae e l'Universo "accelerato" (2003)









"Lo spazio non e' poi cosi' remoto. E' giusto a un'ora di macchina da qui,... se solo la vostra auto potesse scalare il cielo"

F. Hoyle

Credits:

·La musica di "A come Andromeda" e' del Mo. Mario Migliardi

•Fonti iconografiche: Hoyle's commemorative site at www.hoyle.org.uk, Clemson University photo archive (USA) at www.astro.clemson.edu, Emilio Segre' visual archive at photos.aip.org